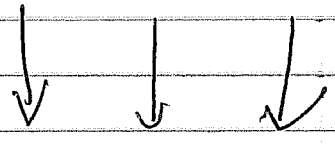
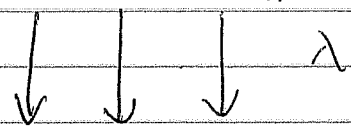


Cross Section Definitions

For 9/26/06
Start ch2 with this class (A)

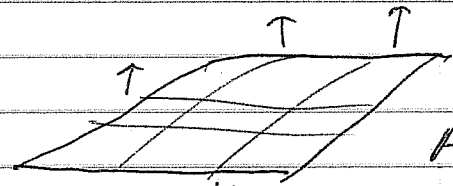
Sunlight $I_0 (W/m^2)$



Black Surface,
Area $A, (m^2)$

Flat
Surfaces,

$$I_0 A \gg \lambda$$



$A = \text{Area}$

$a = \text{albedo}$

$$\text{Absorptance} = (1 - a)$$

Power Absorbed

$$I_0 A$$

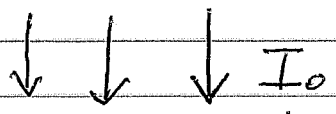
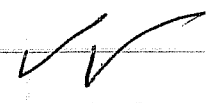
$$\sigma_{\text{abs}} = A$$

Absorption
Cross Section }

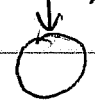
Power Absorbed

$$I_0 A (1 - a)$$

$$\sigma_{\text{abs}} = A (1 - a)$$



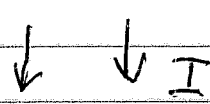
Absorption



— Thing (Particle, molecule, flea, etc)

✓ Absorption indicated by less light thru thing.

Scatter



Power Removed From beam:

①

$$P_{\text{ext}} = I_0 \sigma_{\text{ext}} = I_0 (\sigma_{\text{abs}} + \sigma_{\text{sc}})$$

$$P_{\text{abs}} = I_0 \sigma_{\text{abs}}$$

$$P_{\text{sc}} = I_0 \sigma_{\text{sc}}$$

②

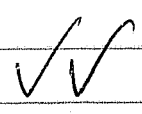
$$\text{Let } w \equiv \frac{\sigma_{\text{sc}}}{\sigma_{\text{ext}}}$$

③

WE Single Scatter Albedo (Compare: Albedo from multiple scatter by many things)

$$P_{\text{abs}} = (1 - w) \sigma_{\text{ext}} I_0$$

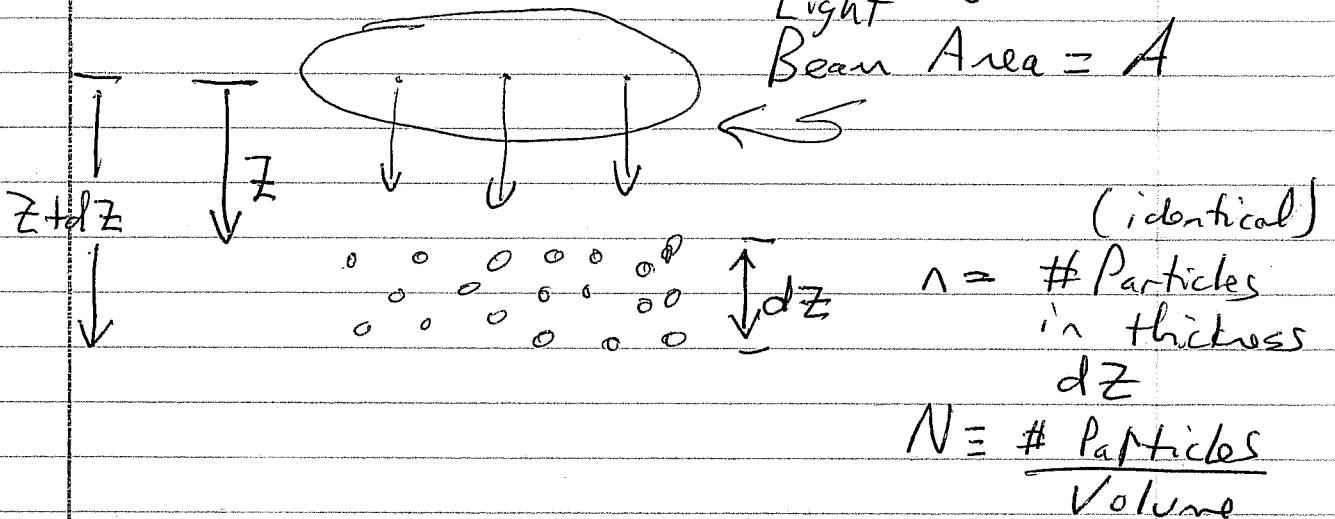
$$P_{\text{sc}} = w \sigma_{\text{ext}} I_0$$



9/26/06
Class
notes

Particles $\rightleftharpoons$ "Things"

Single Scatter Optics of Many Things (B)



Power Removed = $I(z) n \sigma_{ext}$
in dz
or

$(I(z) - I(z+dz)) A = I(z) N A dz \sigma_{ext}$
or

$-dI = I(z) N \sigma_{ext} dz$

$\int_{I_0}^{I(z)} \frac{dI}{I} = - \int_0^z N \sigma_{ext} dz'$

$\sigma_{ext} = \beta_{ext} = \text{extinction coefficient}$

$\ln \frac{I}{I_0} = - N \sigma_{ext} z$

$I(z) = I_0 e^{-N \sigma_{ext} z}$

Students
Like
This

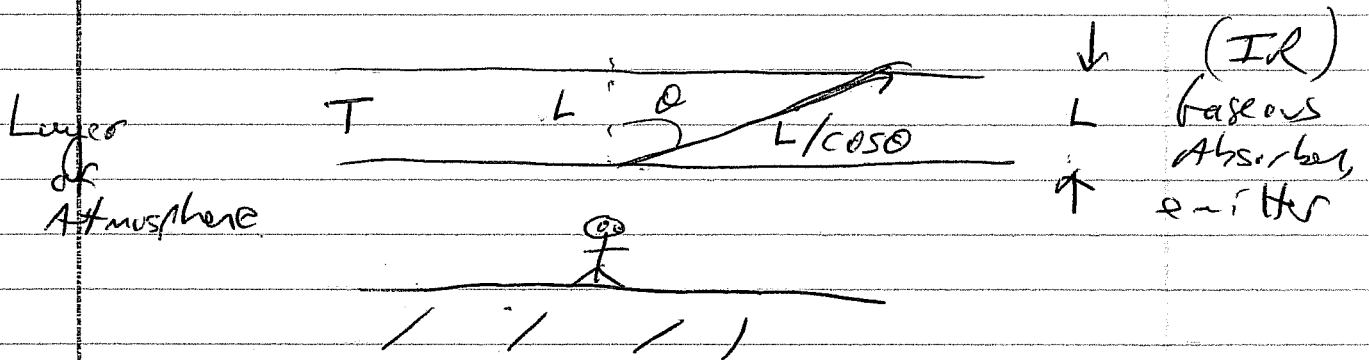
Danger
Beam.

Generally describes ONLY the direct beam attenuation. IF $N \sigma_{scat} \geq 0.5$, multiple scattering can be very important.

2.1A)

CH2 Outline:

- 1) Exponential Attenuation, Skin Depth
[From the Power Point Presentation]
- 2) Directional Emissivity



$$\text{Transmissivity}(\theta) = e^{-\frac{\kappa L}{\cos\theta}}$$

$$\epsilon(\theta) = 1 - e^{-\kappa L/\cos\theta}$$

$\theta \uparrow \cos\theta \downarrow$, see more atmosphere at temperature T ,

$$\epsilon(\theta) \uparrow$$

- 3) Flux Divergence, Largest?

Heating Rate: $\frac{\partial T}{\partial t} = -\frac{1}{\rho_0 c_p} \frac{\partial F}{\partial z}$

Stratosphere

$$F(z)$$

$$F(z) = F_0 e^{-\kappa z}$$

$$-\frac{\partial F}{\partial z} = + \kappa F_0 e^{-\kappa z}$$

Earth

What value of κ gives greatest heating rate?

2.2

$$\frac{\partial}{\partial x} \left(\frac{\partial F}{\partial z} \right) = 0 = F_0 [e^{-xz} - xze^{-xz}] = 0$$

$$\Rightarrow xz = 1 \text{ after factoring out } e^{-xz}$$

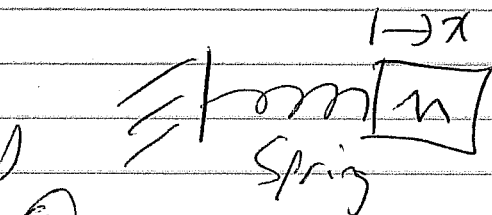
Absorption optical depth $\Rightarrow 1$, max heating rate, trace in stratosphere.

How do we talk about the death of photons?

Go to Pg 2 2.2A, 2.2B

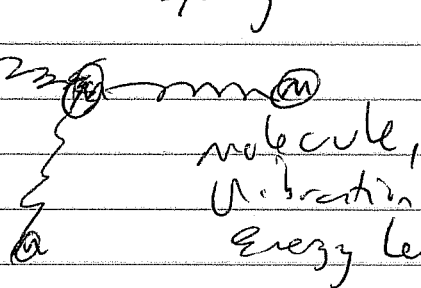
Harmonic Oscillators

All the world is a mass



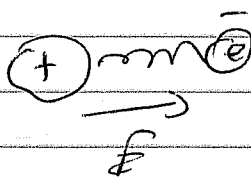
$$F = -kx$$

$$V = \frac{1}{2} kx^2 \text{ Potential Energy}$$



Energy levels

"Resonance" Associated with Energy levels -



Charge drive by external Electric field

$$\vec{E}$$

Nature does All it can to
get things to share their energy

4.65

Ch2

Tropopause

Absorbs IR
Emits little
of atm

$$\text{Slab Emission} = E(\lambda) P_e(\lambda, T)$$

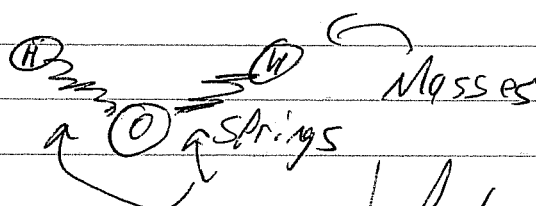
Power
Planck Electromagnetic Energy
function

$P_e(\omega, T)$

/ / / / / Ground
T_g

(1) Molecules have kinetic energy associated with translation - moving around. They move around!

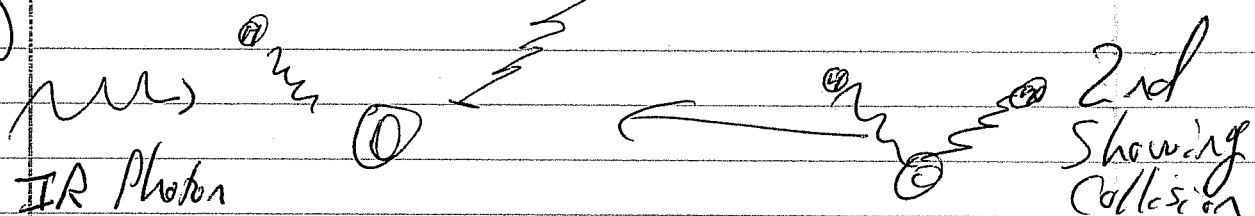
(2) Molecules have "Modes" $\Rightarrow$ (Ways they can dance)



Rotations Moments of Inertia

Absorption

(3)



IR Photon

λ just right to "kick" a mode

Photon energy $\Rightarrow$ "kick"

kick shared with others during collision.

Result: heat How Shared?

Heat Capacity: $C_p = R + \frac{3}{2}R + \frac{2R}{2} + \frac{?R}{2}$

vibration
translation rotation

4.66

KE

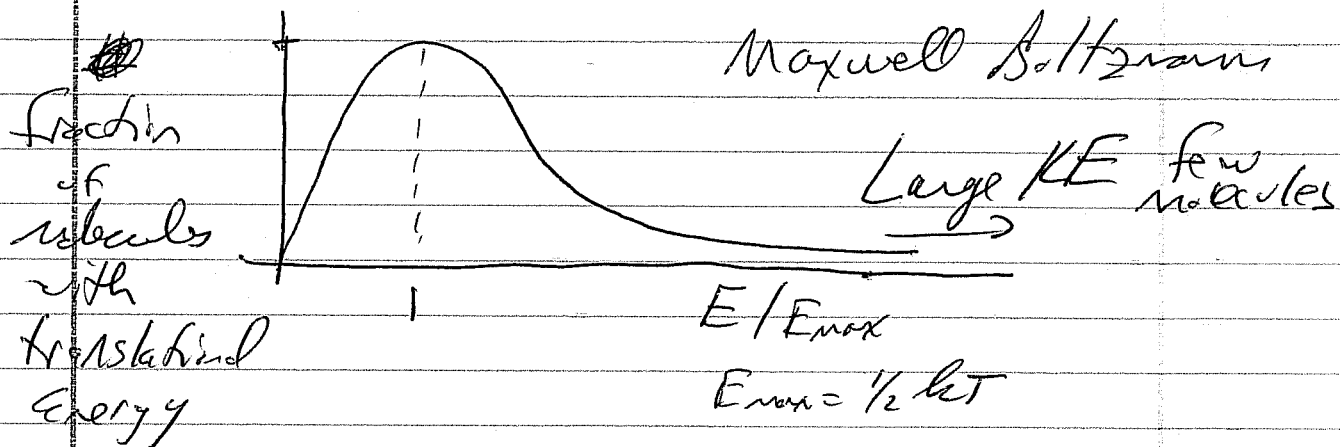
(4) ~~Q~~ Is the translational energy enough to excite all modes?

Is the tickle strong enough to excite the modes?

translation ✓

rotation ✓

Vibration not so much



(5) Run the Film in Reverse:

